

# **Dynamics of the plasma sheet in the magnetotail: interrelation of turbulent flows and thin current sheet structures**

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# Main topics: both important and still disputable

- Intense nonlinear disturbances in the **MAGNETOTAIL PLASMA SHEET**.

Most remarkable features are

(a) medium scale plasma **FLOWS**  
and magnetic field variations:

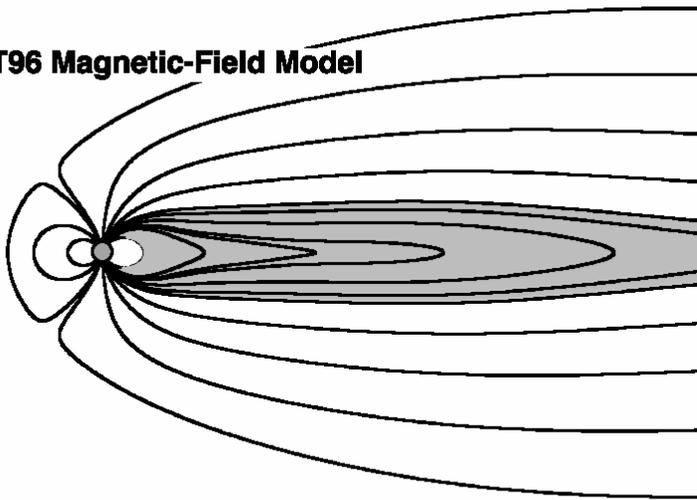
**MHD TURBULENCE**

(b) Sporadic thin **CURRENT SHEETS**  
embedded in the plasma sheet.

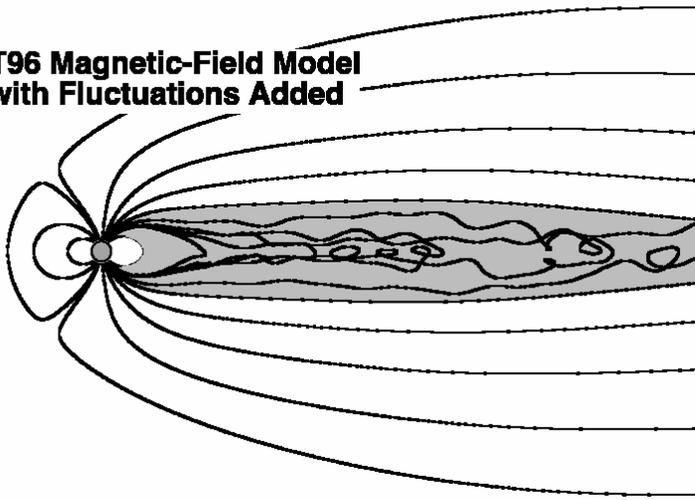
**WHAT IS THE INTER-RELATION?**

- **Dynamical nature of thin current sheets.**  
**SPONTANEOUS FORMATION OF NONLINEAR KINETIC STRUCTURE**
- **How do thin current sheets relate to MAGNETIC RECONNECTION?**
- **Range of temporal and spatial scales.**  
**Why and how is SLOW EVOLUTION of configuration sporadically interrupted by FAST LOSSES OF EQUILIBRIUM?**

**T96 Magnetic-Field Model**



**T96 Magnetic-Field Model  
with Fluctuations Added**

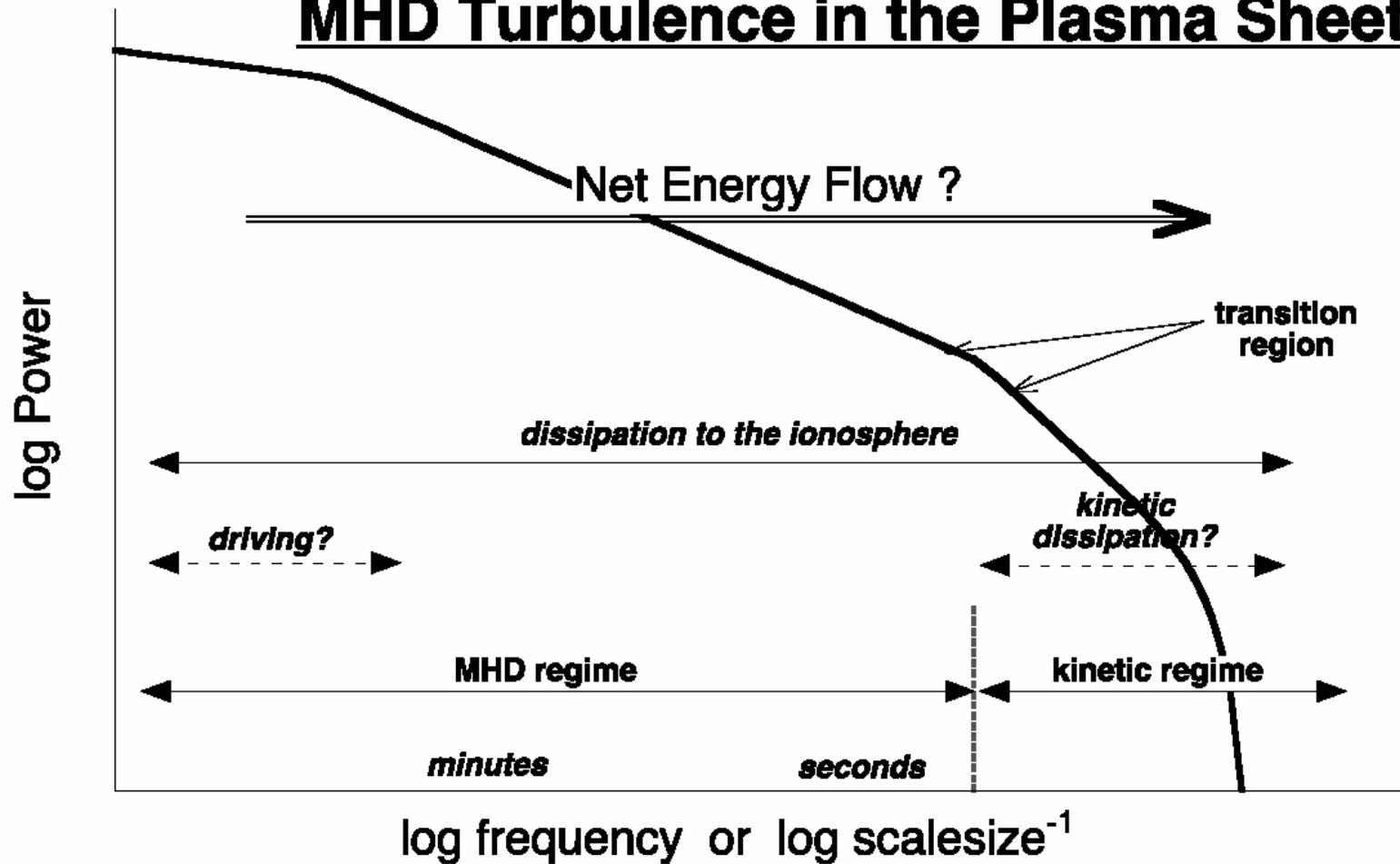


*Borovsky et al, 2003*

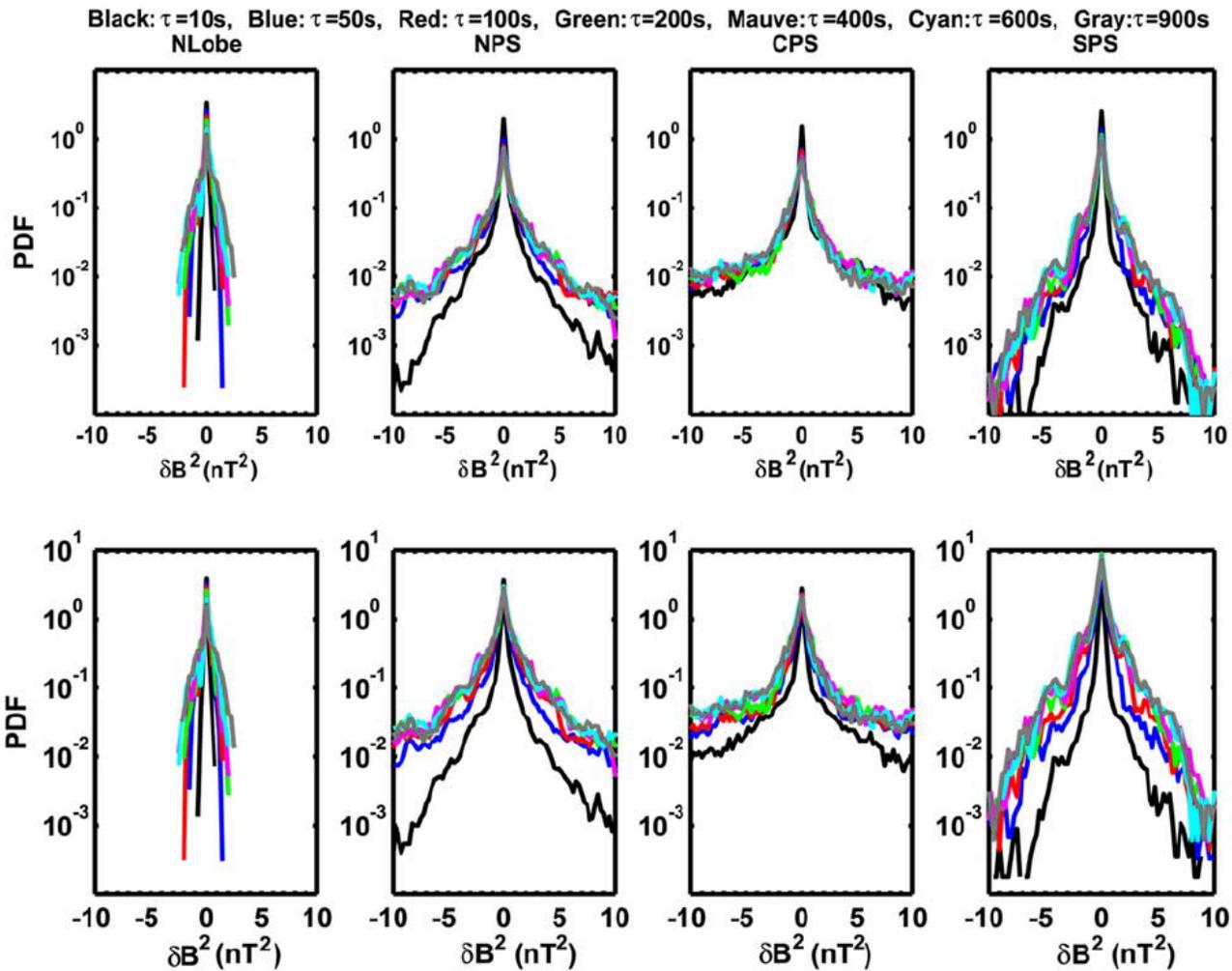
**Complicated pattern of convective motions in the plasma sheet of the geomagnetic tail associated with magnetic field variations occurring on MEDIUM SCALES, large as compared to the plasma sheet thickness and small as compared to global dimensions. These motions and magnetic variations are interpreted as manifestations of specific, basically TWO-DIMENSIONAL MHD TURBULENCE.**

*(Borovsky et al, Antonova et al)*

# MHD Turbulence in the Plasma Sheet

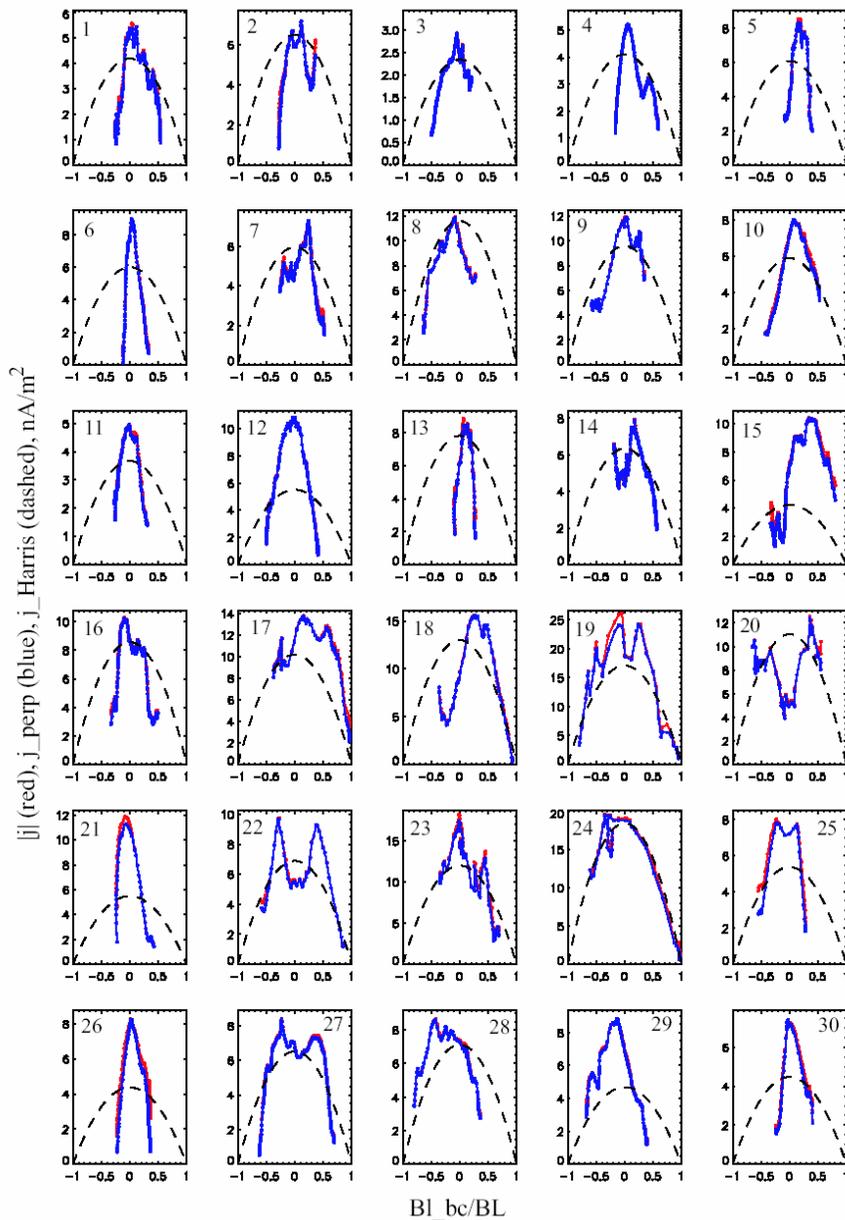


*Borovsky et al, 2003*



*Weigand et al 2005*

**Non-Gaussian probability density functions**



|j| (red), j\_perp (blue), j\_Harris (dashed), nA/m<sup>2</sup>

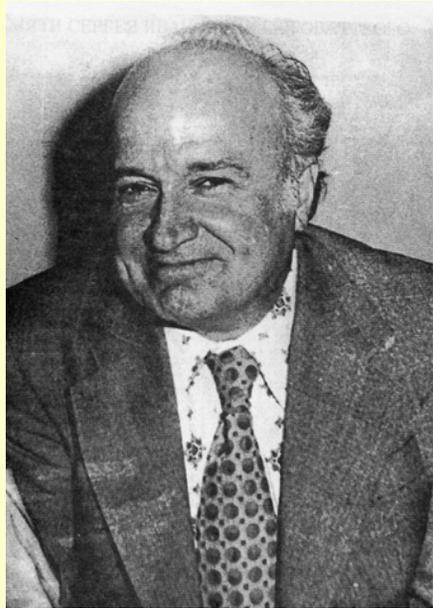
Bl\_bc/BL

## Thin current sheets in the geomagnetic tail

Hodograms of the current density  $\mathbf{j} = \mu_0^{-1} \nabla \times \mathbf{B}$  absolute value (blue) and perpendicular component

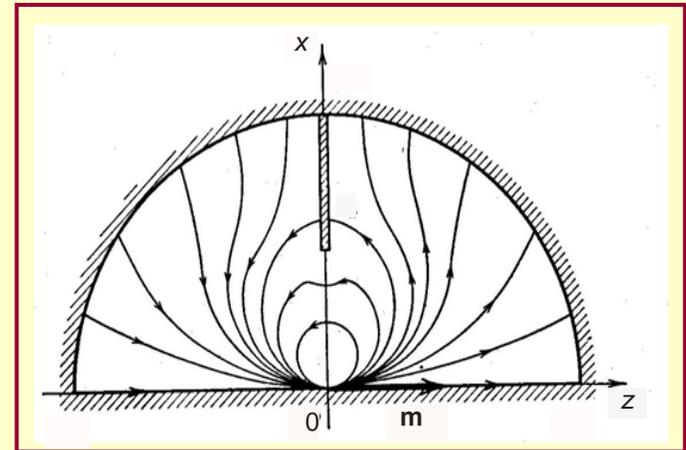
$$j_{\perp} = \sqrt{j_m^2 + j_n^2} \quad (\text{red})$$

versus main magnetic field ( $B_l$ ). Dashed lines show the Harris function



Sergey Ivanovich Syrovatsky

A simple two-dimensional model of the magnetosphere (Syrovatsky and Somov, 1974).



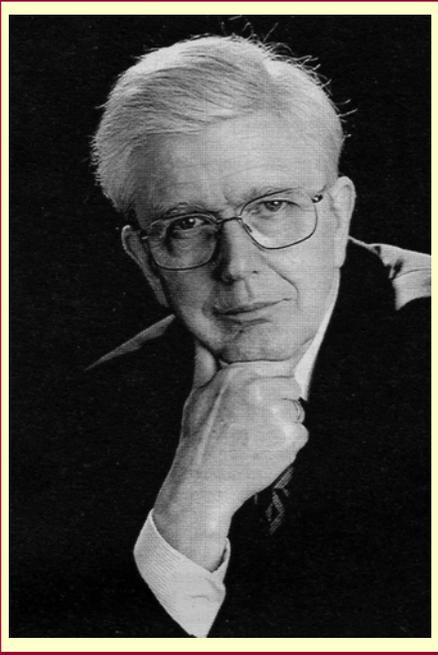
The essence of that modeling was that in a plasma with  $\beta \ll 1$  all the **CURRENTS ARE CONCENTRATED IN THIN CURRENT SHEETS (CS)**, and outside them the magnetic field is curl-free. In two dimensions, we thus simply obtain Laplace equation  $\Delta A = 0$  for the only nonzero component of the vector-potential, with proper conditions on the boundaries.

**OUTSIDE THE CS**, the characteristic time scale is the Alfvénic time scale

$$\tau \approx l / V_A \approx l \sqrt{4\pi\rho} / B$$

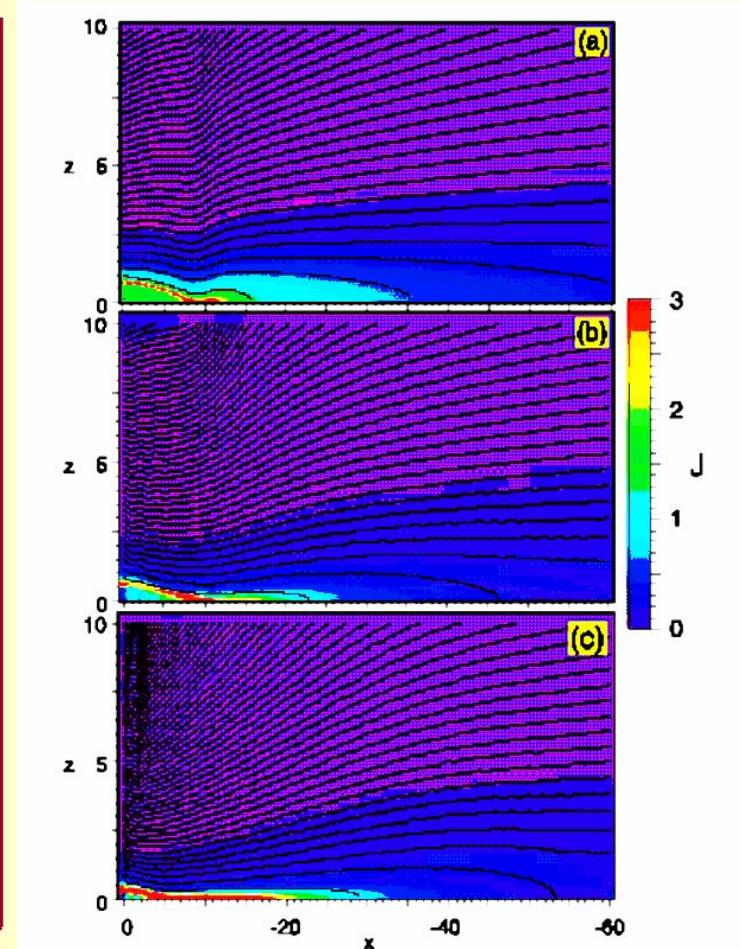
Reconfiguration occurs “instantly”, on a small time scale  $T_1$ , with low density  $\rho$  and high  $B$ . **THE MODEL IS APPLICABLE!**

# Critical role played by thin CS



Karl Schindler

- Theory and simulation: **thin CS inevitably form** in reconfiguration occurring in MHD systems.
- At that point, **MHD becomes no more valid**: a singularity appears in the current density distribution.
- Those singularities are **ONE-DIMENSIONAL: CURRENT SHEETS**
- This greatly simplifies the problem of **KINETIC** approach becoming necessary at that stage.
- Role of the CS: energy transformation,  $\mathbf{jE} > 0$ , over a large portion of the CS



Magnetic field lines and color-coded current density for near-critical states

(a) **quasi-static** theory; (b) MHD simulation with the same boundary deformation as in (a), but with increased amplitude; (c) MHD simulation with tailward **propagating perturbation**.

(J. Birn, K. Schindler, and M. Hesse, *J. Geophys. Res.*, 108(A9), 2003)

**MHD equilibrium equation:**  $\frac{1}{4\pi} \mathbf{B} \cdot \nabla \mathbf{B} = \nabla \left( p + \frac{B^2}{8\pi} \right)$

**Two-dimensional Grad-Shafranov model:**  $\mathbf{B} = \nabla A \times \nabla y \quad p = p(A)$

**Schindler's asymptotic solution,  $l_z / l_x \ll 1$ :**

$$p + \frac{B^2}{8\pi} = \hat{p}(x) \quad z(x, A) = \int_{A_0(x)}^A \frac{dA}{\sqrt{8\pi [\hat{p}(x) - p(A)]}}$$

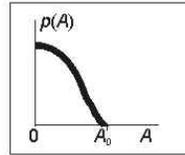
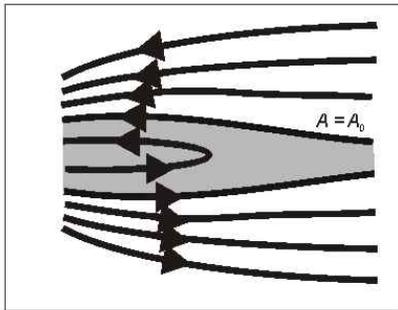
**Tail outer boundary:**  $a(x) = \int_{A_0(x)}^{A_b} \frac{dA}{\sqrt{8\pi [\hat{p}(x) - p(A)]}}$

**Quasi-static evolution is adiabatic:**  $S = pV^\gamma$  and entropy are conserved

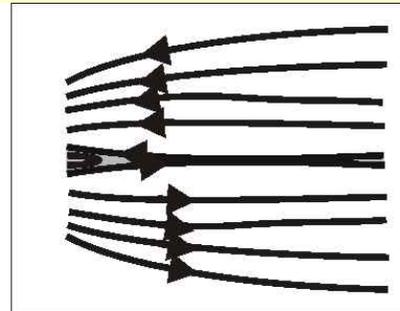
**DENSITY IS CONCENTRATED IN THE CENTRAL PLASMA SHEET**

**RECONFIGURATION IN THE CENTRAL PLASMA SHEET IS SLOW,  
ON A SCALE  $T_2$ ,  $T_2 \ll T_1$**

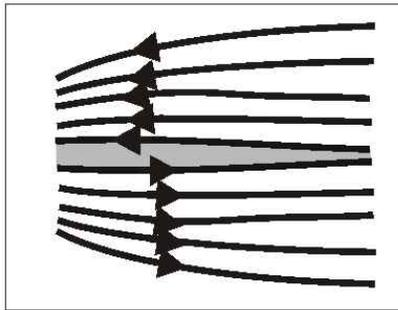
**LOSS OF EQUILIBRIUM IS "INSTANTANEOUS"!**



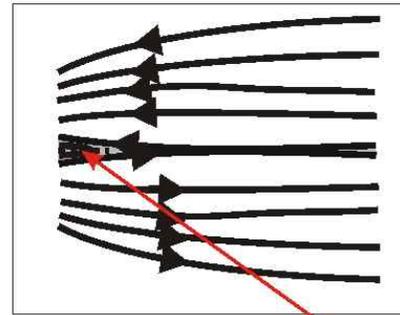
**(a) initial configuration with a thick plasma sheet**



**(c) non-equilibrium configuration with a thin current sheet is formed after a quick reconfiguration initiated by nonlinear tearing mode generation**



**(b) quasi-static adiabatic evolution leads to equilibrium with a thin current sheet**



**(d) locally a new equilibrium is formed, with an embedded anisotropic kinetic current sheet**

# Simulation: basic features

- Initially: hot plasma in a Harris-type CS
- Cold uniform plasma background

$$B_n \neq 0; B_n \ll B_0.$$

- 1D hybrid code: ions treated as macroparticles, electrons as cold massless background; self-consistent electromagnetic fields.
- Simulation domain 6 times larger than the CS thickness.
- Spatial cell size =  $0.085\lambda_0$
- Time step =  $0.1/\Omega_0$
- About 150000 macroparticles
- Normalization parameters:

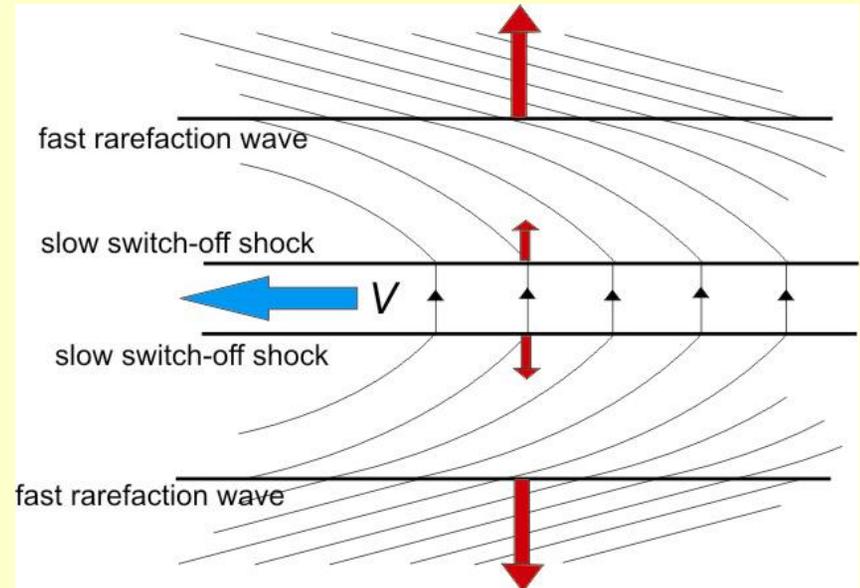
$$B_0 = B_t(z = \infty, t = 0)$$

$$E_0 = B_n V_A / c = B_n B_t / c (4\pi m_i N_0)^{1/2}$$

$$N_0 = N^{(h)}(z = 0, t = 0)$$

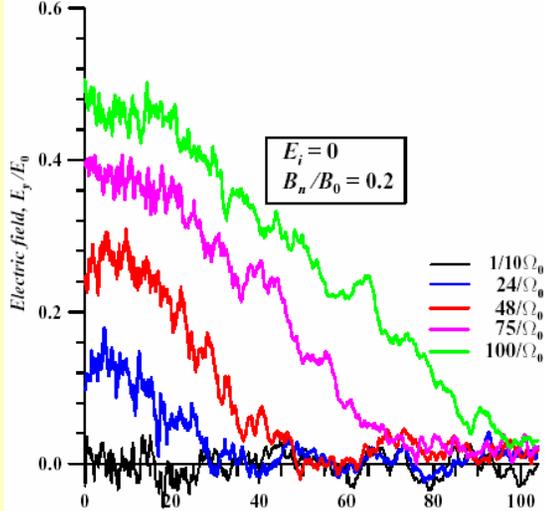
$$N^{(c)} = N_0$$

## MHD prototype



decay of a non-equilibrium discontinuity with field reversal and  $B_n \neq 0$

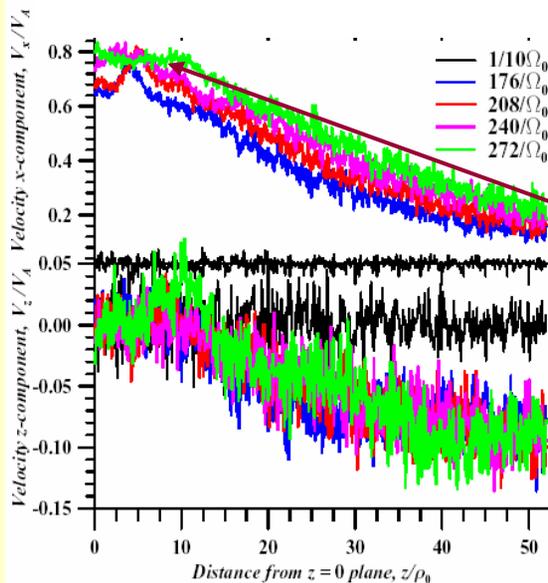
# Time evolution of the current sheet: embedded extremely thin shock structure is formed



Initial stages: fast MHD wave ahead of slow “switch-off” shock

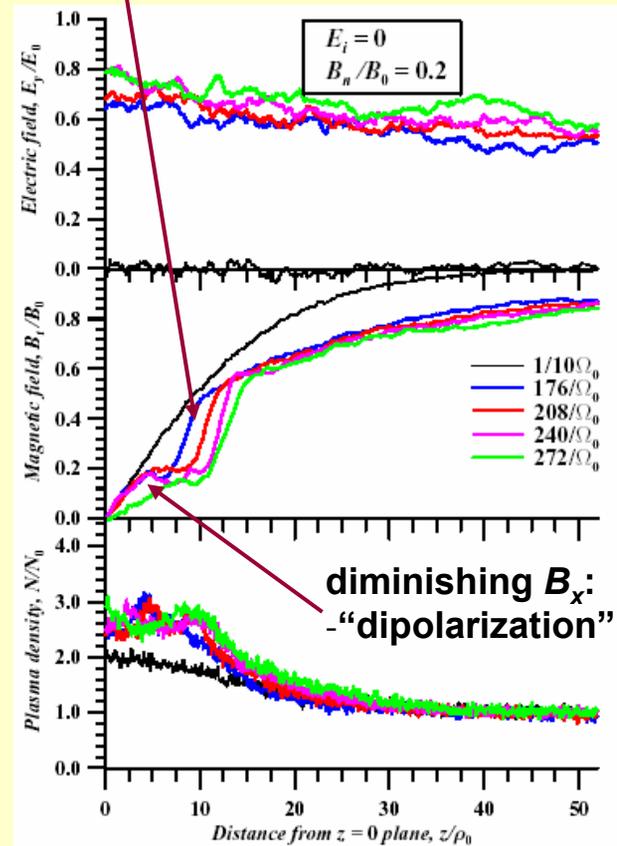
$$B_n / B_0 = 0.2$$

Later stages



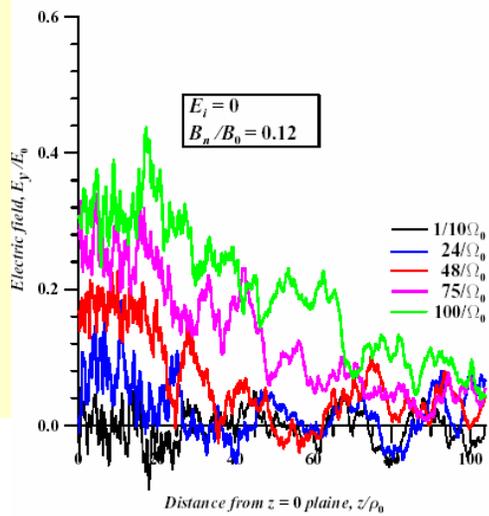
fast ( $v \sim V_A$ ) convection along x at the center

convection towards CS between the fast and slow fronts



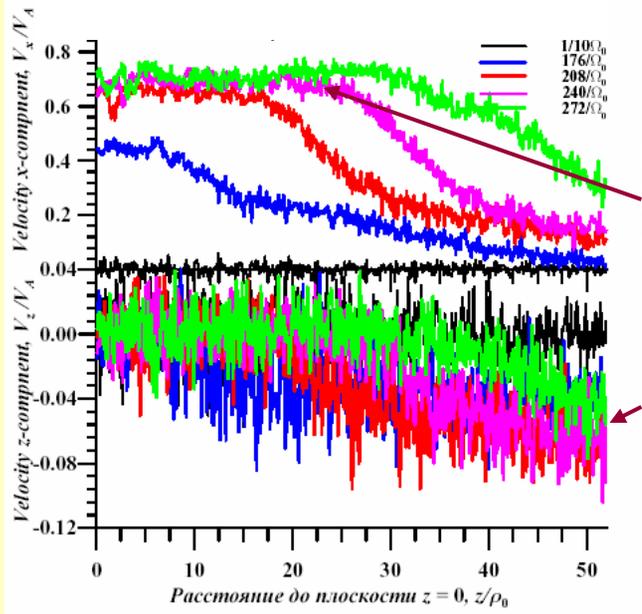
diminishing  $B_x$ : -“dipolarization”

# Time evolution of the current sheet: embedded extremely thin CS is formed – anisotropic Forced Current Sheet



Initial stages: fast MHD wave ahead of slow disturbance

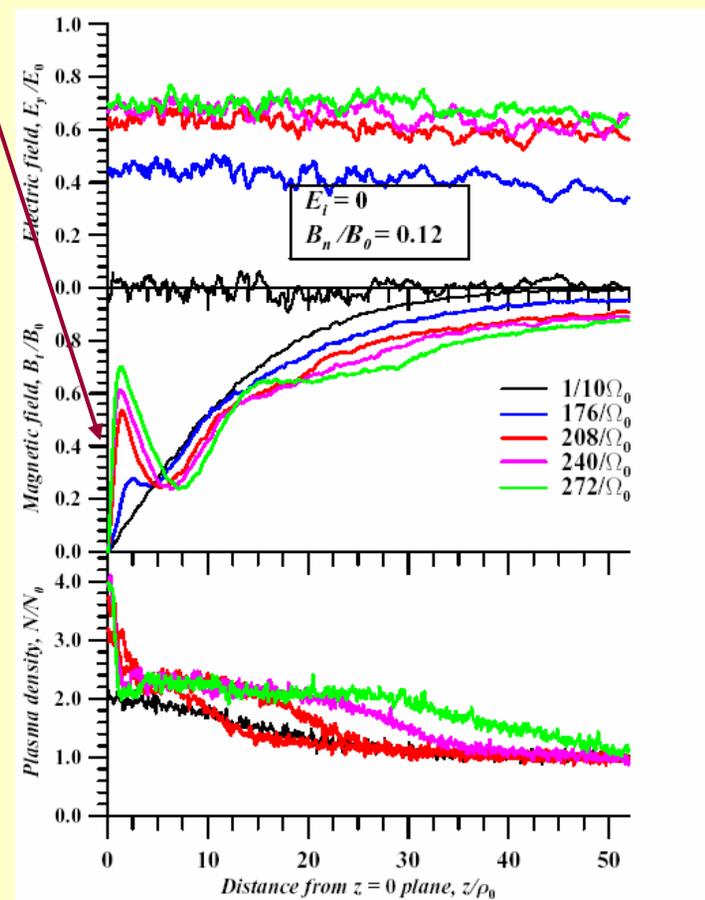
$B_n/B_0 = 0.12$

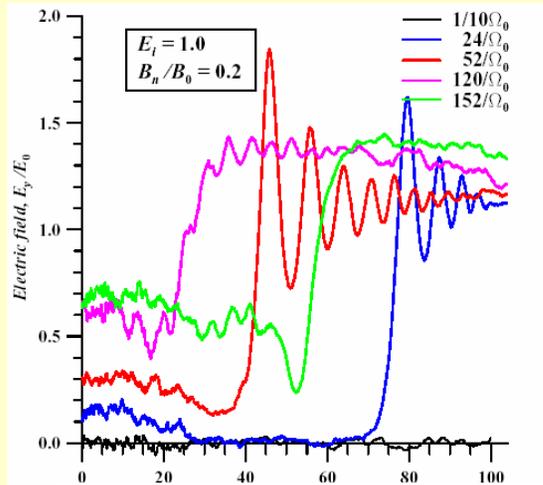


Later stages

Bulk velocity near  $V_A$  in the double-flow region

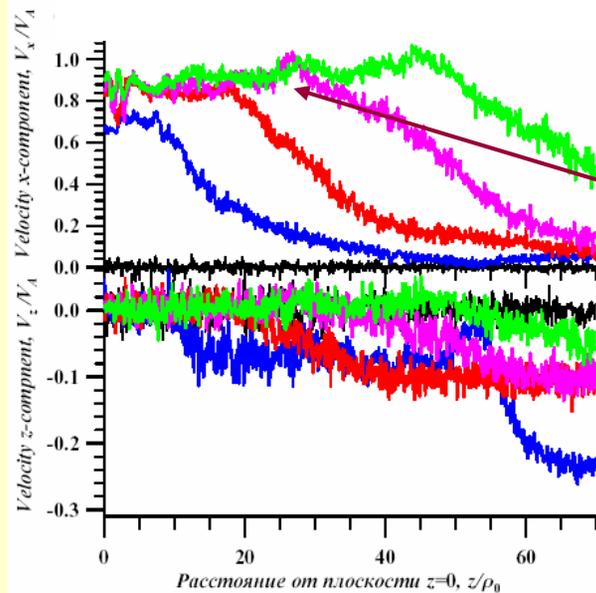
convection towards CS outside the double-flow region





**Initial stages: fast collisionless shocks, incident and reflected**

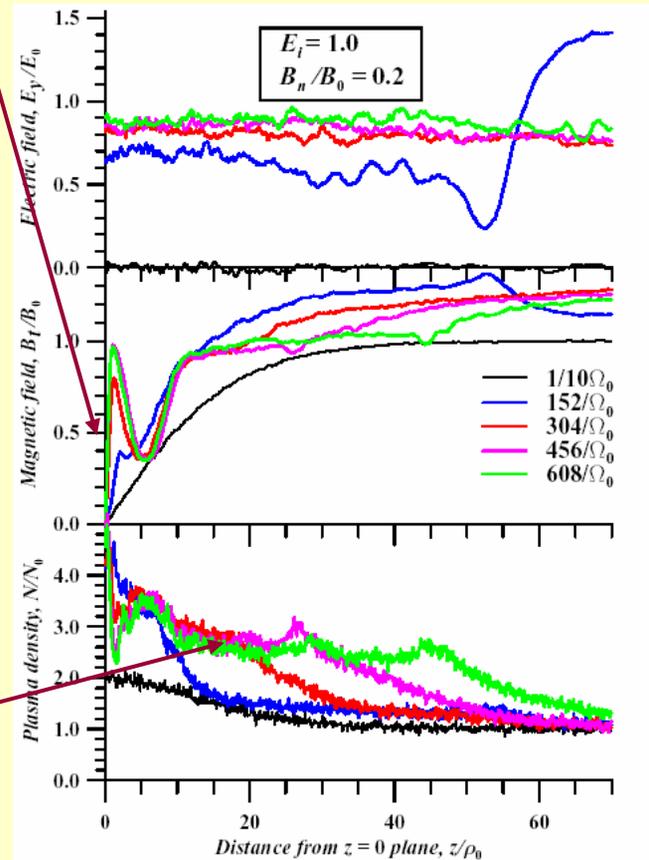
**Time evolution of the current sheet, with an external trigger: embedded extremely thin CS is formed – anisotropic Forced Current Sheet**

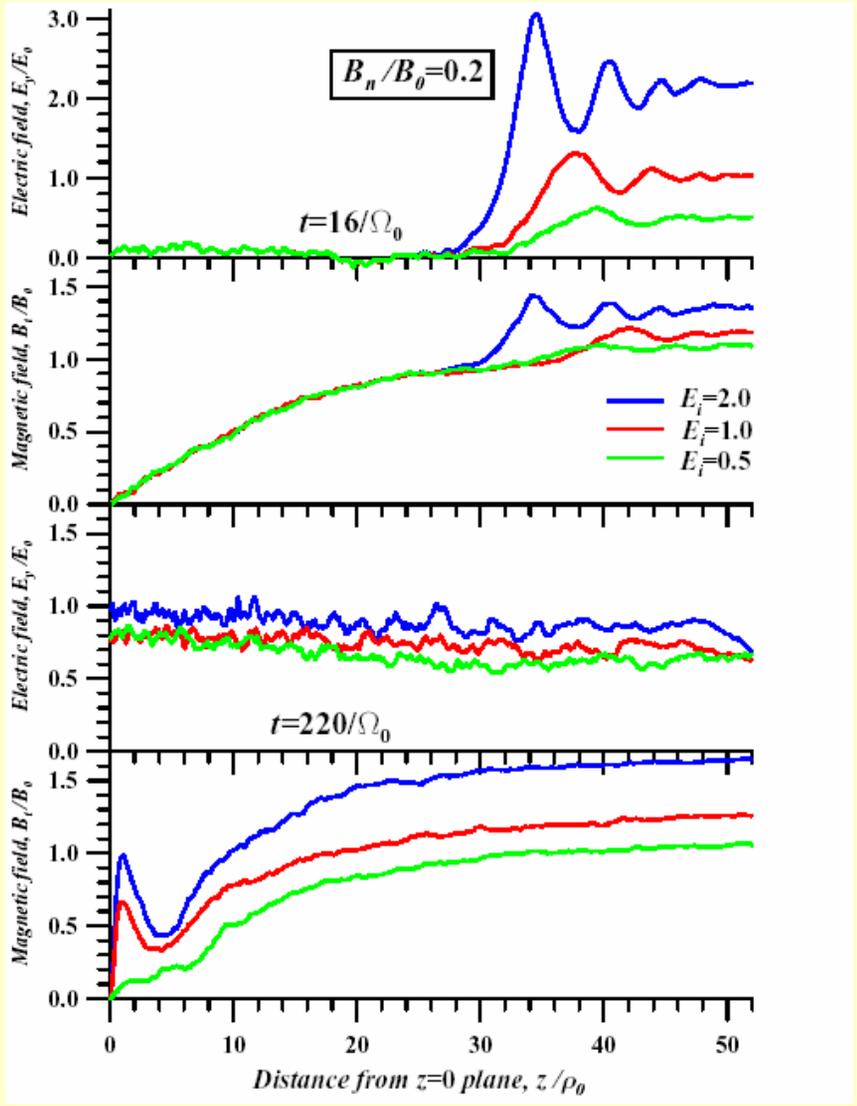
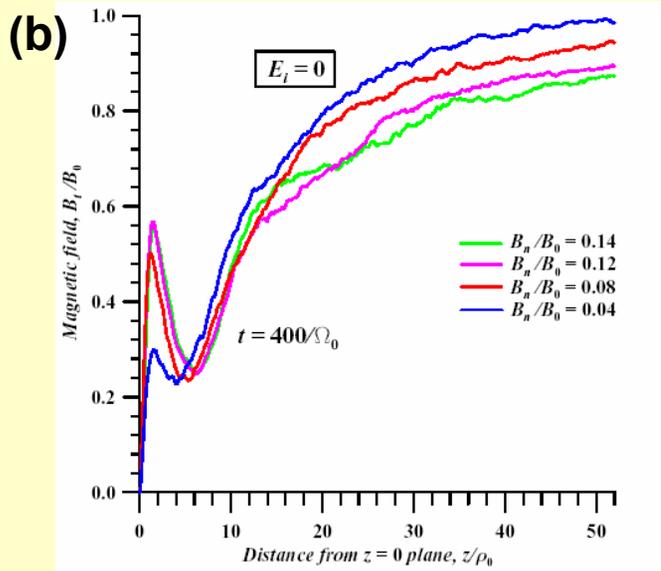
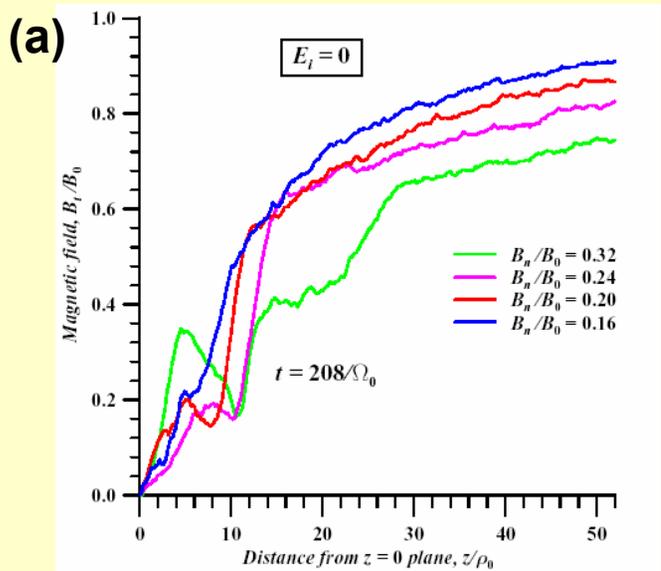


**Later stages**

**Bulk velocity near  $V_A$  in the double-flow region**

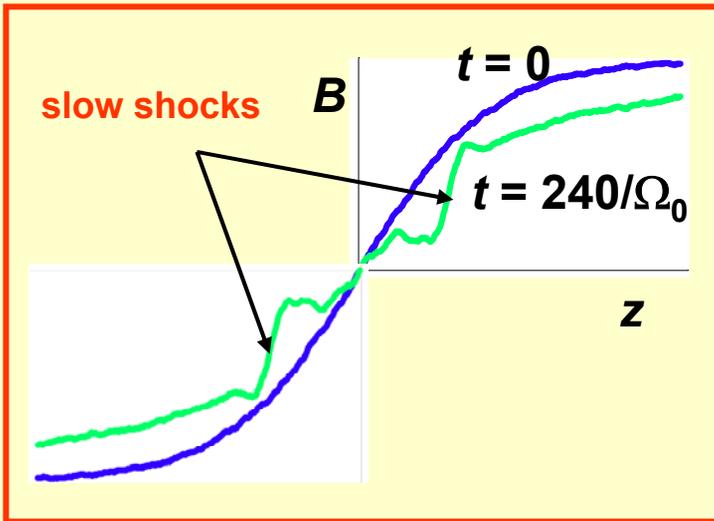
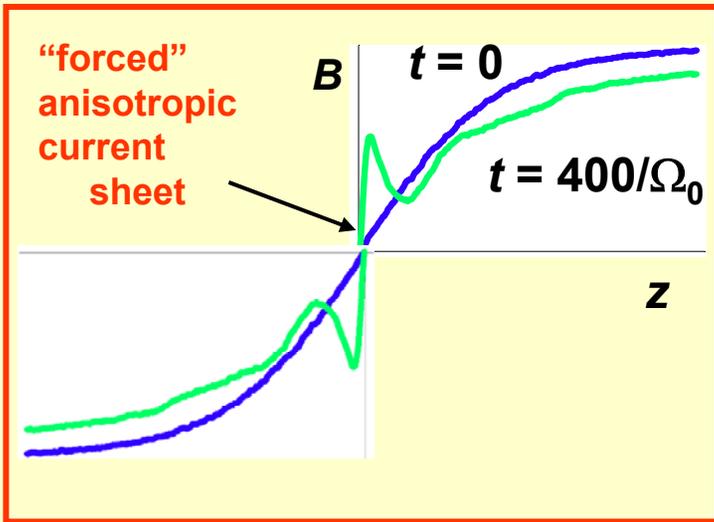
**Density is nearly doubled in the double-flow region**



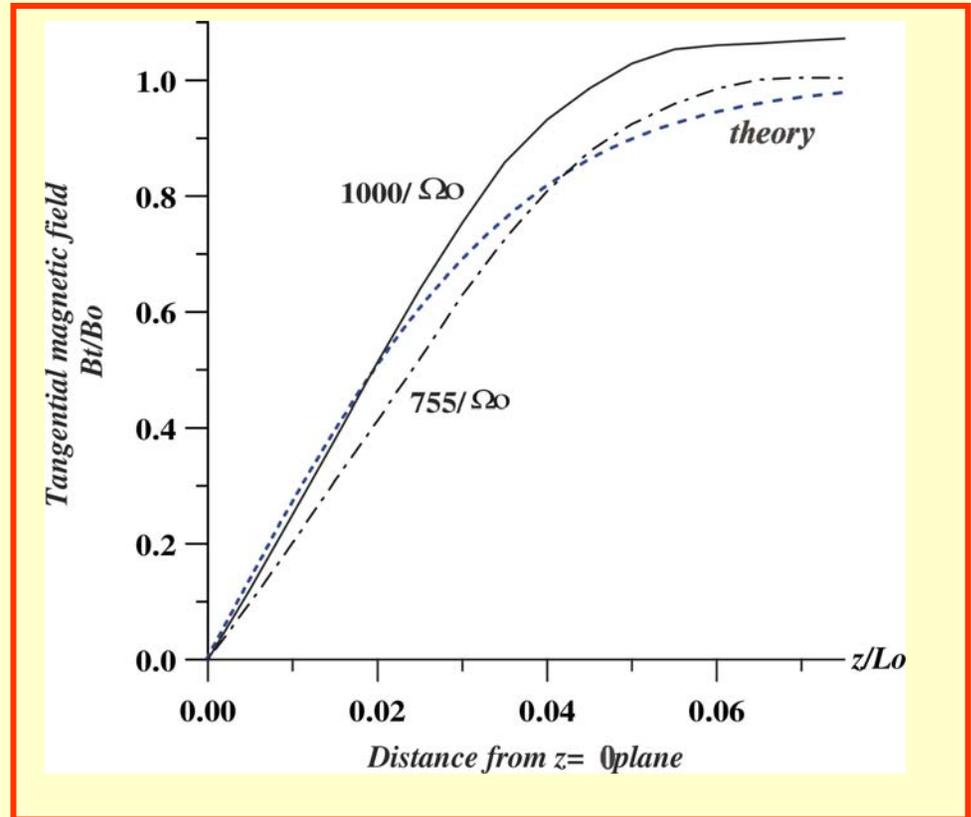


Formation of either a shock (a) or anisotropic FCS (b) depending on  $B_n/B_0$  ratio. Critical  $B_n/B_0 = 0.14$

Formation of anisotropic FCS depending on trigger intensity



**Formation of thin embedded current sheets**

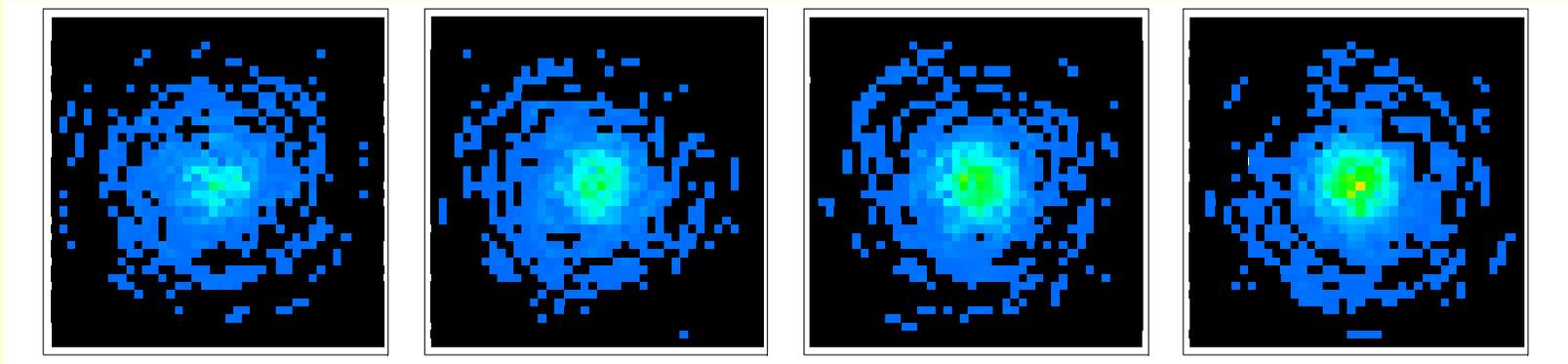


### Consistency with theory

- $B(z)$  profile
- finite energy transformation rate:
- electric field  $E$  and the Poynting vector

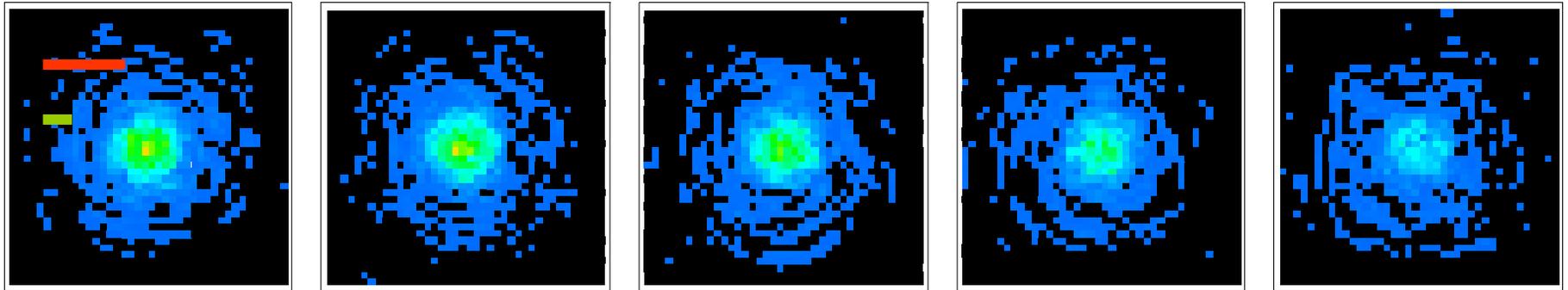
$$P = c \frac{EB_t}{4\pi}$$

$$= \frac{B_t^2 B_n}{(4\pi)^{3/2} (N_0 m_i)^{1/2}}$$



$V_z = -2.45v_T^{(c)}$

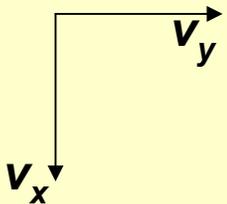
$V_z$  increase  $\rightarrow$



$V_z = 0$

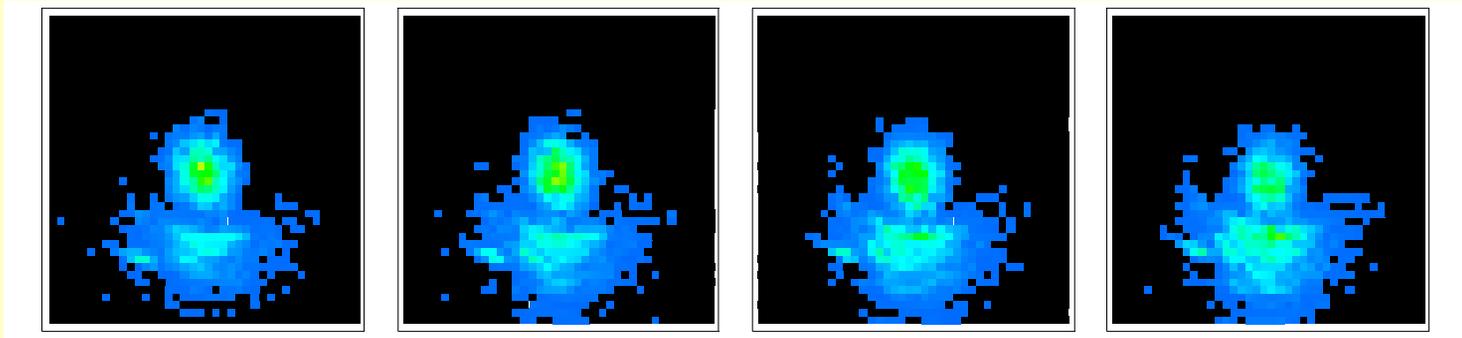
$V_z$  increase  $\rightarrow$

$V_z = 2.45v_T^{(c)}$



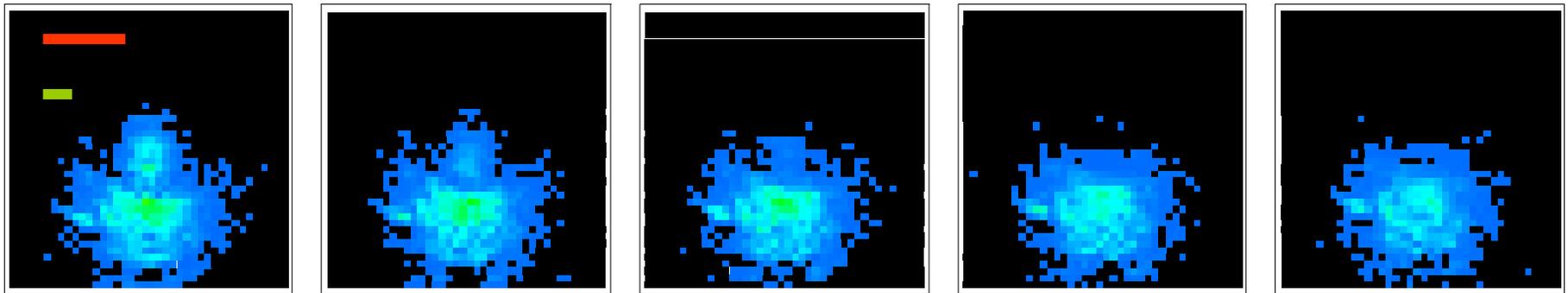
Successive cross-sections of the velocity distribution  
 No trigger;  $B_n/B_0 = 0.2$ ;  
 $t = 0, Z = 0$

—  $v_T^{(h)}$   
 —  $v_T^{(c)}$



$V_z = -2.45v_T^{(c)}$

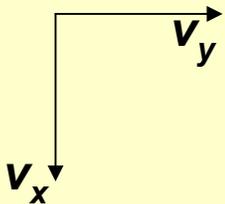
$V_z$  increase  $\rightarrow$



$V_z = 0$

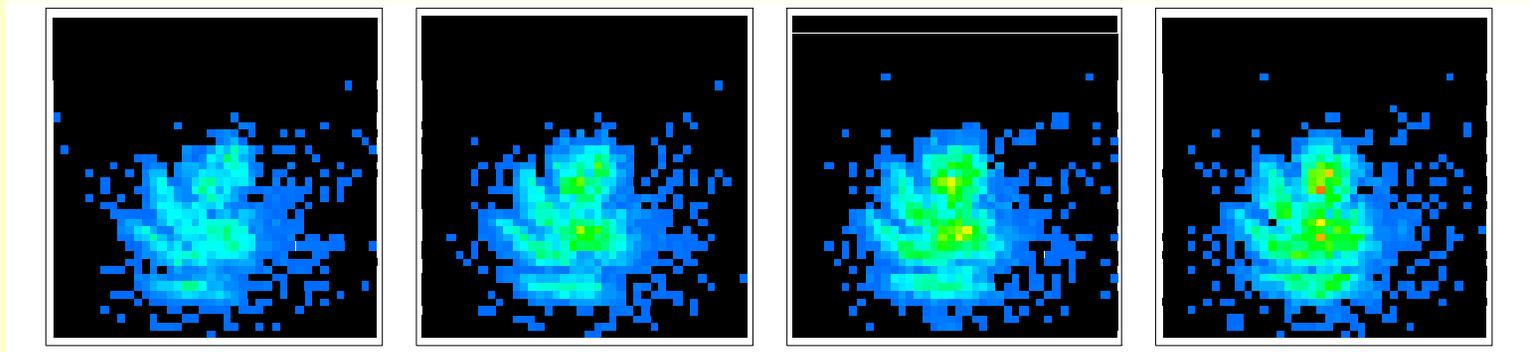
$V_z$  increase  $\rightarrow$

$V_z = 2.45v_T^{(c)}$



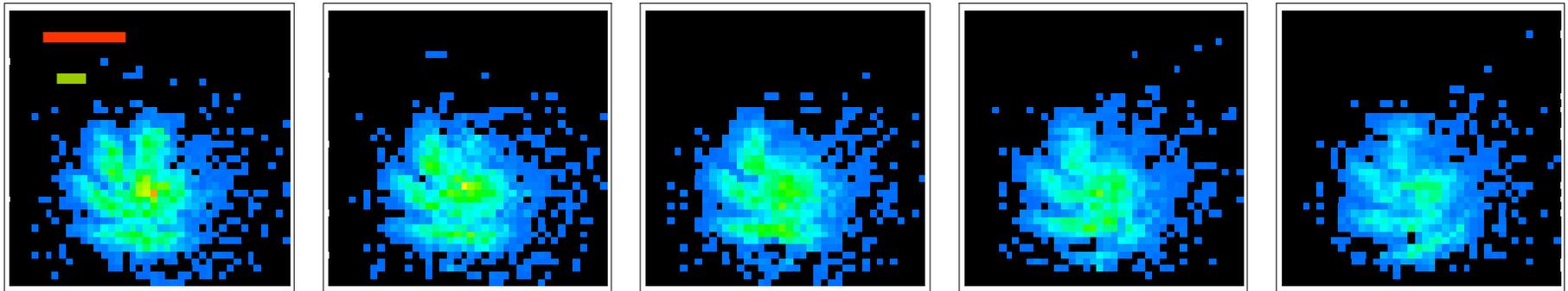
Successive cross-sections of the velocity distribution  
 No trigger;  $B_n/B_0 = 0.2$ ;  
 $t = 4/\Omega_0$ ,  $Z = 12.2\rho_0$

—  $v_T^{(h)}$   
 —  $v_T^{(c)}$



$V_z = -2.45v_T^{(c)}$

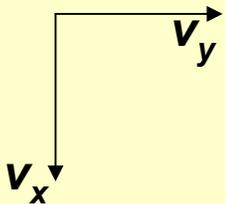
$V_z$  increase  $\rightarrow$



$V_z = 0$

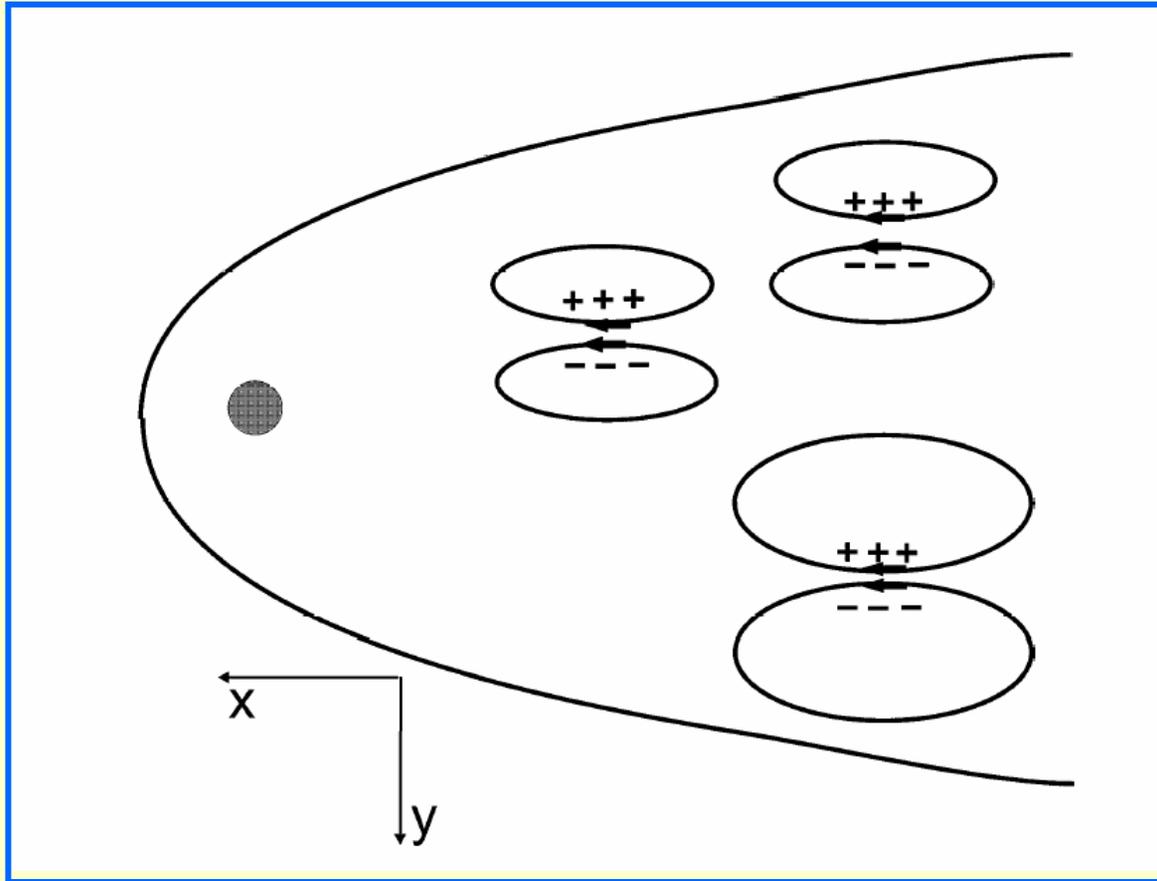
$V_z$  increase  $\rightarrow$

$V_z = 2.45v_T^{(c)}$



Successive cross-sections of the velocity distribution  
 No trigger;  $B_n/B_0 = 0.12$ ;  
 $t = 4/\Omega_0$ ,  $Z = 0$

—  $v_T^{(h)}$   
 —  $v_T^{(c)}$



**Patches of magnetic field merging in the geomagnetic tail.  
Sporadic electric fields and plasma flows**

## CONCLUSIONS

- Thin CS initially produced by flux-conserving (MHD) motions, is sporadically affected by **LOSS OF EQUILIBRIUM**. Via fast MHD disturbance, the loss of equilibrium is remotely **INDUCED BY A NONLINEAR (TEARING, BALLOONING) INSTABILITY MECHANISM**.
- A short time scale  $T_1$  is associated with the MHD disturbance propagating in the tail lobes, with their relatively strong magnetic field and low plasma density.
- On those areas of CS where the induced loss of equilibrium occurs, nonlinear quasi-one-dimensional evolution is started. Its crucial feature is **SELF-ORGANIZATION** expressed in spontaneous formation of **EMBEDDED EXTREMELY THIN CURRENT STRUCTURES**. A much longer time scale  $T_2$  is associated with that process in the plasma sheet, with its small normal component of the magnetic field and a greater density.
- In **SIMULATION**, both **SLOW SHOCKS AND ANISOTROPIC “FORCED” CS** are identified as such extremely thin current structures.

## CONCLUSIONS (ctd)

- **KINETIC EFFECTS** provide various features of **ANISOTROPY**: (a) shifted nearly Maxwellian distributions: **FAST BULK FLOWS**; (b) **FIELD-ALIGNED DOUBLE FLOWS**, etc.
- The thin kinetic structures are responsible for large-scale magnetic **MERGING**, i.e. transformation of magnetic energy into energy of plasma flows, occurring over the structure. Together with scattering induced by plasma wave turbulence, this provides a mechanism of **DISSIPATION IN THE ENTIRE SYSTEM**.
- In the magnetosphere, this is the basis for **INTERMITTENT**, sporadic **MAGNETOTAIL ACTIVATIONS** (fast nonlinear tearing or ballooning instability events initiate the whole process).
- The **FAST FLOWS** may **DRIVE TURBULENCE** on shorter spatial scales. In their turn, these motions may serve as an origin for neutral line generation, and **RECONNECTION**. The latter generates signals in MHD modes, propagating in the magnetotail lobes, and thus new fast disturbances on the short time scale  $T_1$  are produced. **INTERMITTENCY IS TYPICAL**.